

# Gravitational Wave Probes of Axion Kinaton

Raymond Co



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Brookhaven Forum November 4<sup>th</sup> 2021

Based on:

1910.02080 RC, Keisuke Harigaya

Phys. Rev. Lett. 124, 111602 (2020)

1910.14152 RC, Lawrence Hall, Keisuke Harigaya

Phys. Rev. Lett. 124, 251802 (2020)

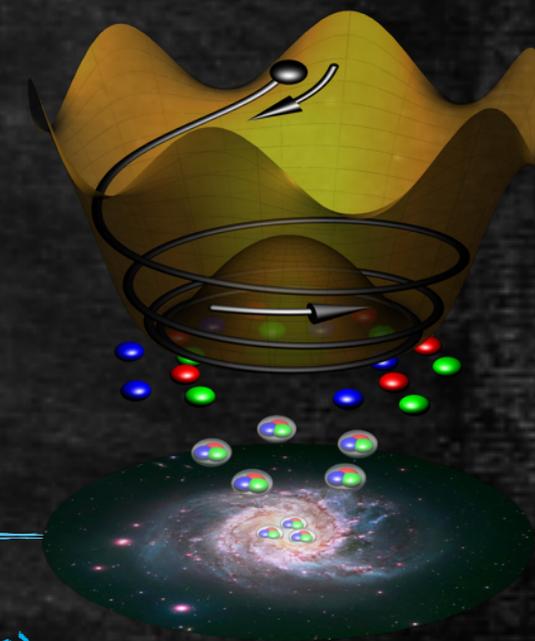
2006.04809 RC, Lawrence Hall, Keisuke Harigaya

JHEP 01 (2021) 172

2006.05687 RC, Nicolas Fernandez, Akshay Ghalsasi, Lawrence Hall, Keisuke Harigaya

JHEP 03 (2021) 017

2108.09299 RC, David Dunskey, Nicolas Fernandez, Akshay Ghalsasi, Lawrence Hall, Keisuke Harigaya, Jessie Shelton





Today

Nobel Prize in Physics 1989

for the invention of the separated oscillatory fields method  
and its use in the hydrogen maser and other atomic clocks

Norman F. Ramsey

died on

November 4<sup>th</sup> 2011

# Early Universe Dynamics

## Axion

### (0) Misalignment mechanism

Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983

### (1) Parametric resonance

RC, L. Hall, K. Harigaya 2017    K. Harigaya, J. Leedom 2019

### (2) - Kinetic misalignment mechanism

RC, L. Hall, K. Harigaya 2019    + K. Olive, S. Verner 2020

### - Axiogenesis

RC, K. Harigaya 2019

### - ALPogenesis

RC, L. Hall, K. Harigaya 2020

### - Lepto-Axiogenesis

RC, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya 2020

### - Tachyonic instability

RC, K. Harigaya, A. Pierce 2020



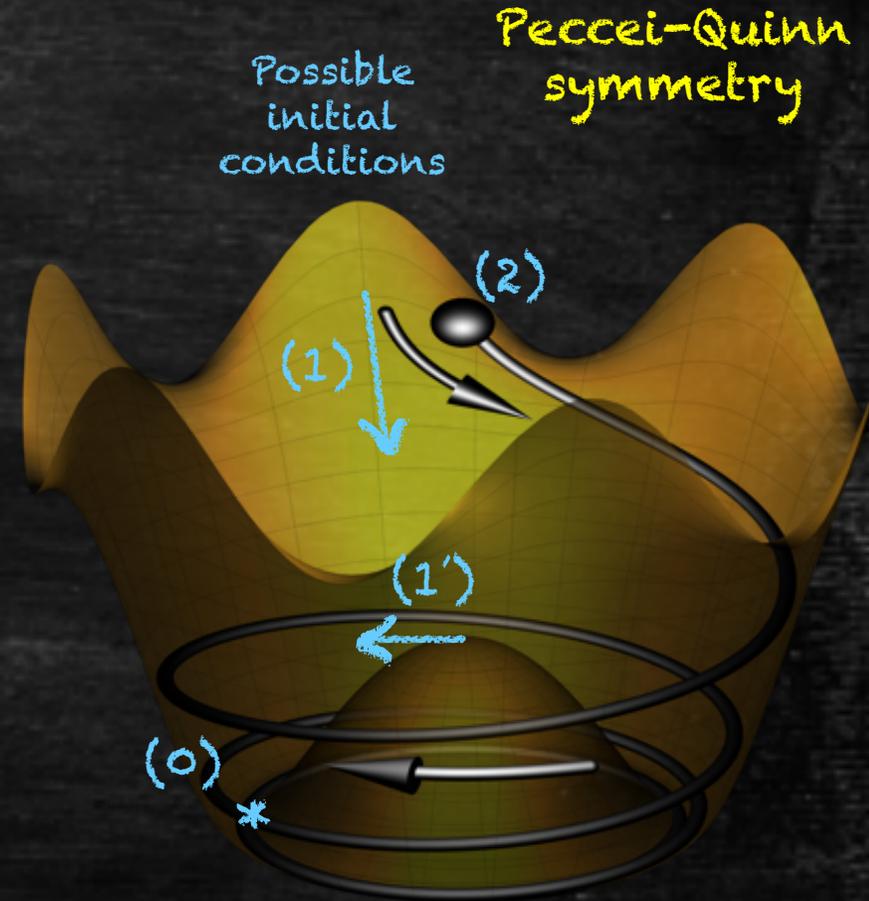
Dark Matter



Baryon Asymmetry



Gravitational Waves



# Axion Rotations

## Axion

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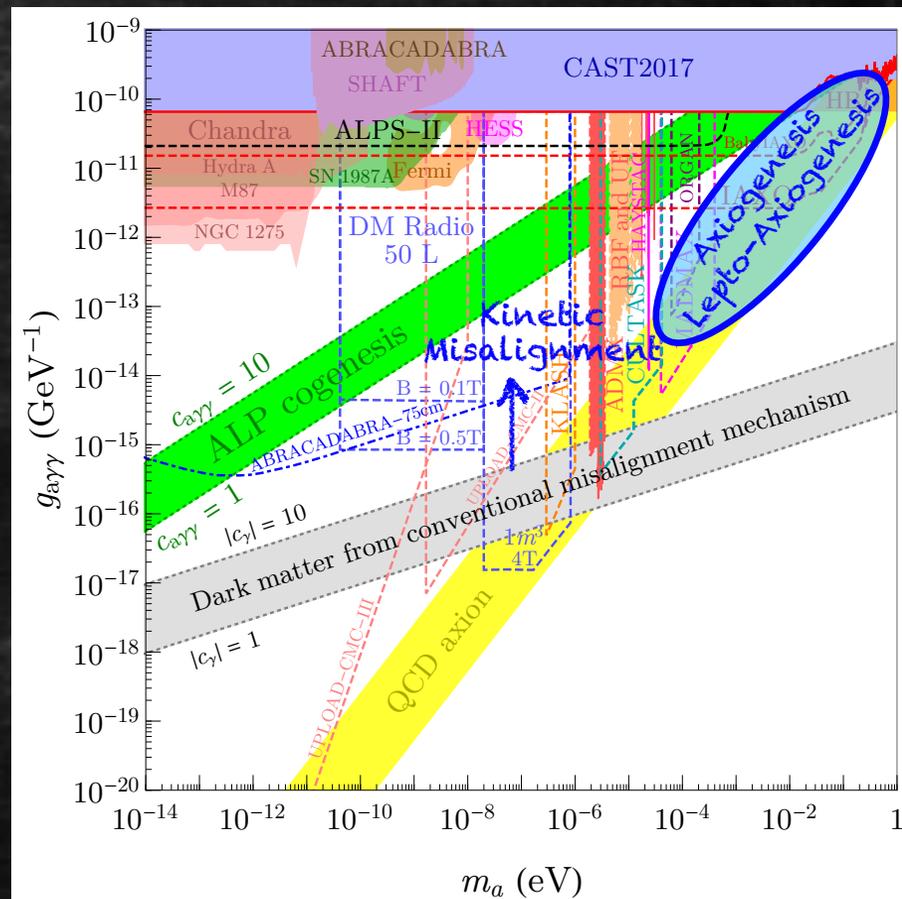
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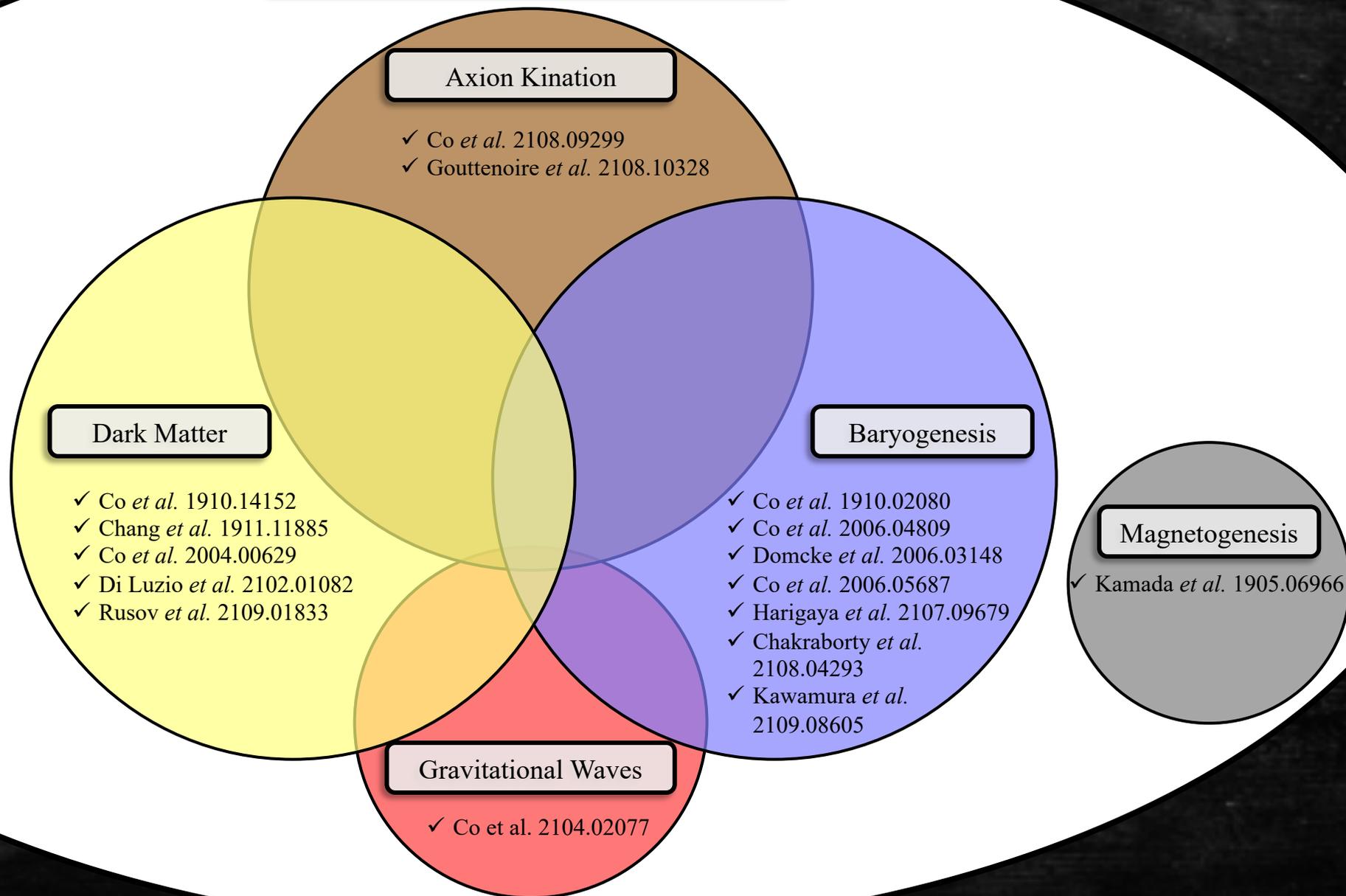
} Dark Matter

} Baryon Asymmetry

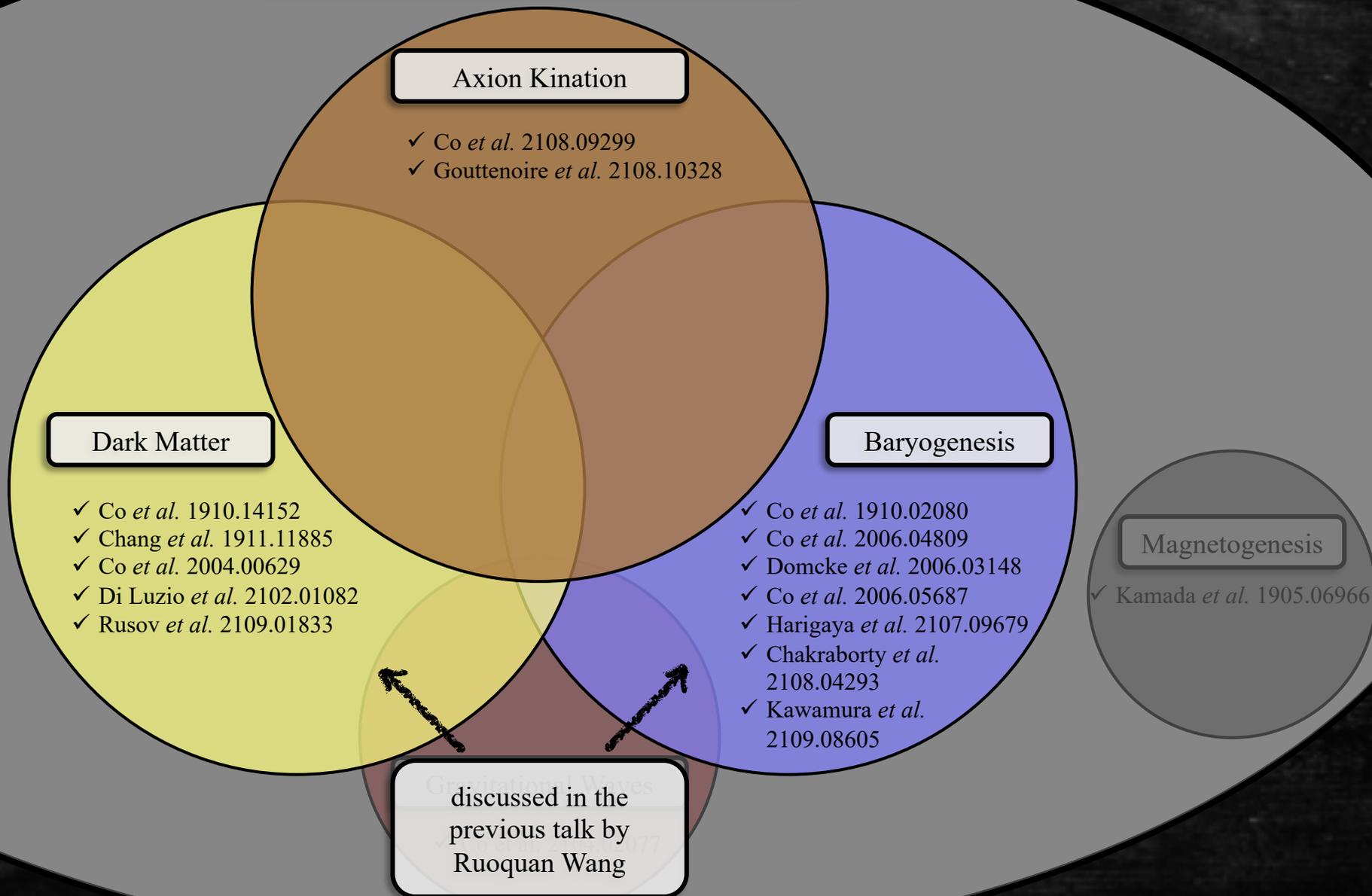
} Gravitational Waves



# Axion Rotations



# Axion Rotations



# Axion Rotations

## Axion Kination

- ✓ Co *et al.* 2108.09299
- ✓ Gouttenoire *et al.* 2108.10328

## Dark Matter

- ✓ Co *et al.* 1910.14152
- ✓ Chang *et al.* 1911.11885
- ✓ Co *et al.* 2004.00629
- ✓ Di Luzio *et al.* 2102.01082
- ✓ Rusov *et al.* 2109.01833

## Baryogenesis

- ✓ Co *et al.* 1910.02080
- ✓ Co *et al.* 2006.04809
- ✓ Domcke *et al.* 2006.03148
- ✓ Co *et al.* 2006.05687
- ✓ Harigaya *et al.* 2107.09679
- ✓ Chakraborty *et al.* 2108.04293
- ✓ Kawamura *et al.* 2109.08605

## Gravitational Waves

- ✓ Co *et al.* 2104.02077

## Magnetogenesis

- ✓ Kamada *et al.* 1905.06966

# What is kination?

9306008 Boris Spokoiny

9606223 Michael Joyce

"One simple alternative - domination by the energy in a kinetic mode of a scalar field which scales as  $1/R^6$ ."

$$\rho_\phi = \frac{1}{2} \left( \dot{\phi}^2 + m_\phi^2 \phi^2 \right) \simeq \frac{1}{2} \dot{\phi}^2$$

Equation of state

$$w = \frac{p}{\rho} = \frac{K - V}{K + V} \simeq 1$$

Evolution

$$\rho_\phi \propto R^{-3(1+w)} = R^{-6}$$

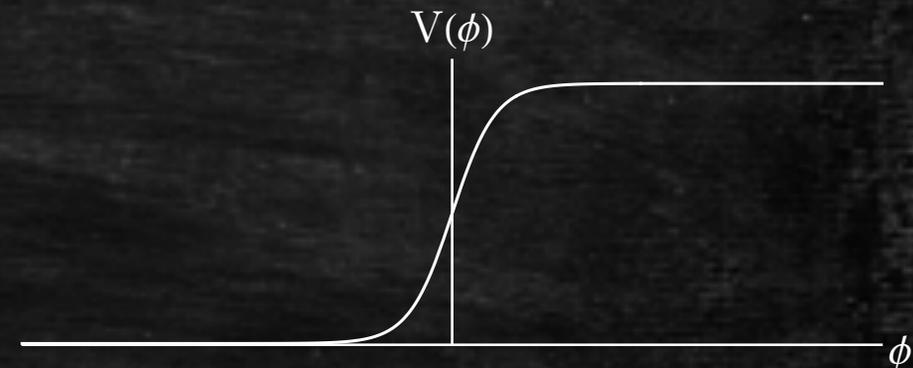
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An unnatural example

Equation of state

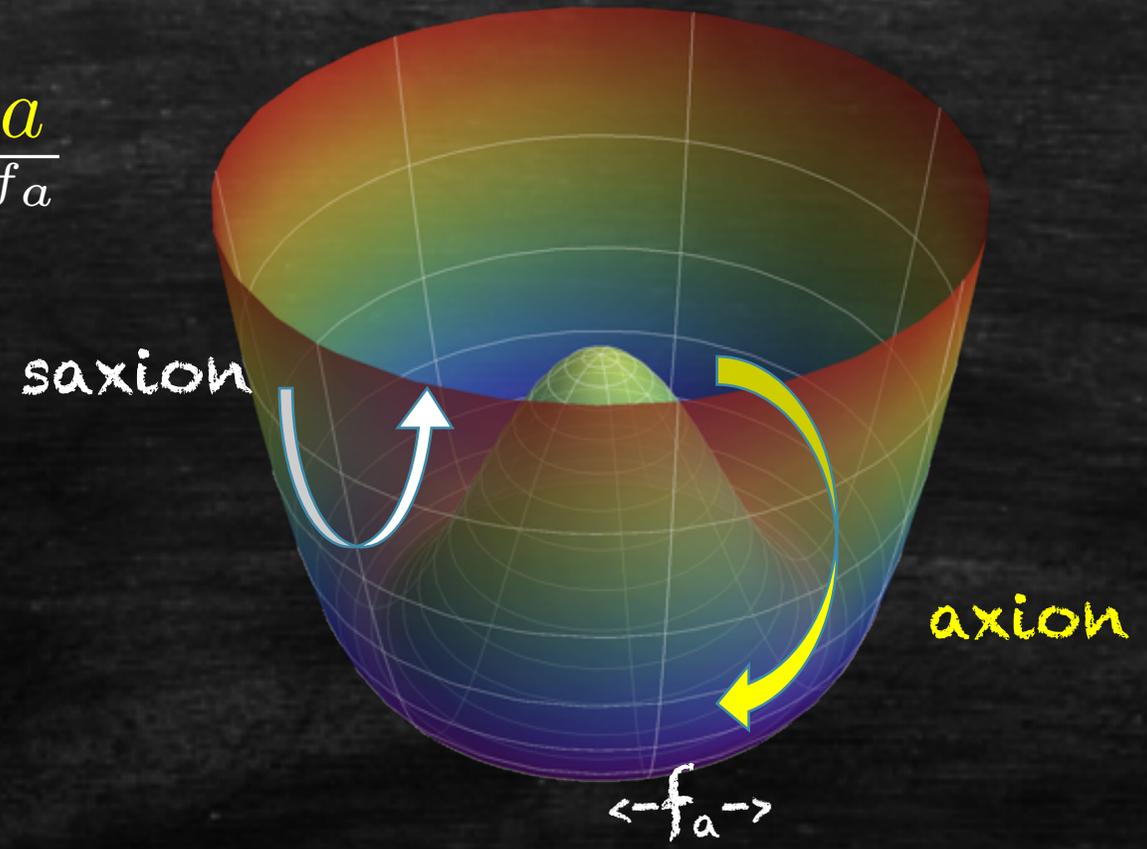
$$w = \frac{p}{\rho} = \frac{K - V}{K + V} \simeq 1$$

Evolution

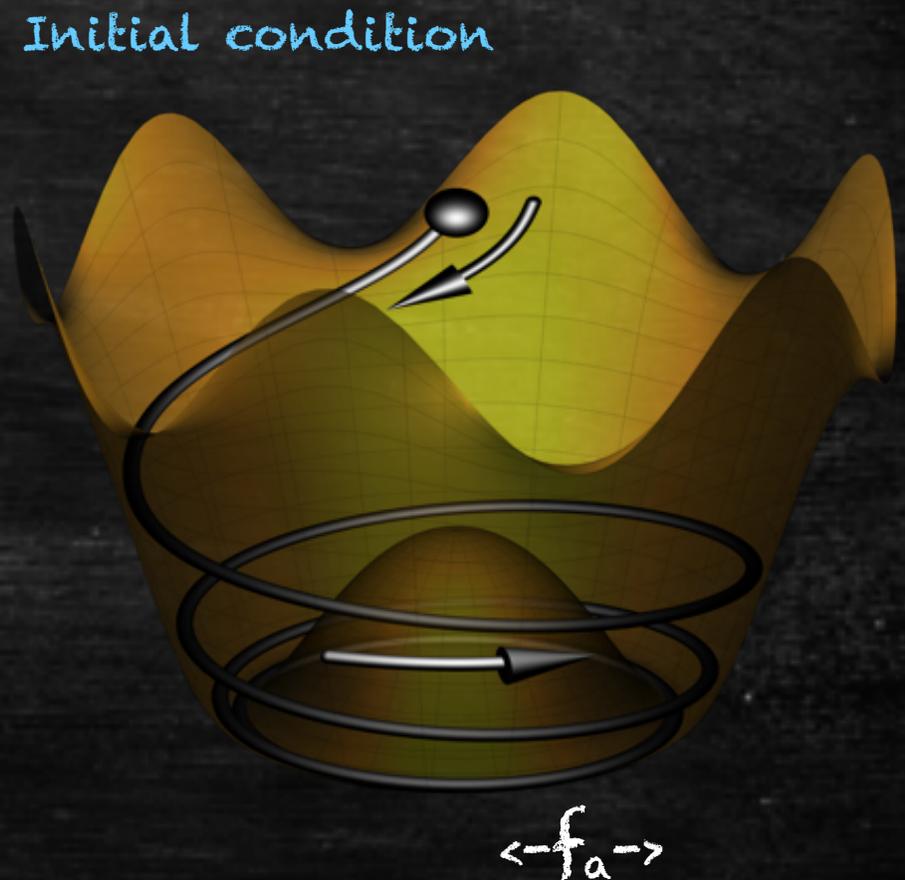
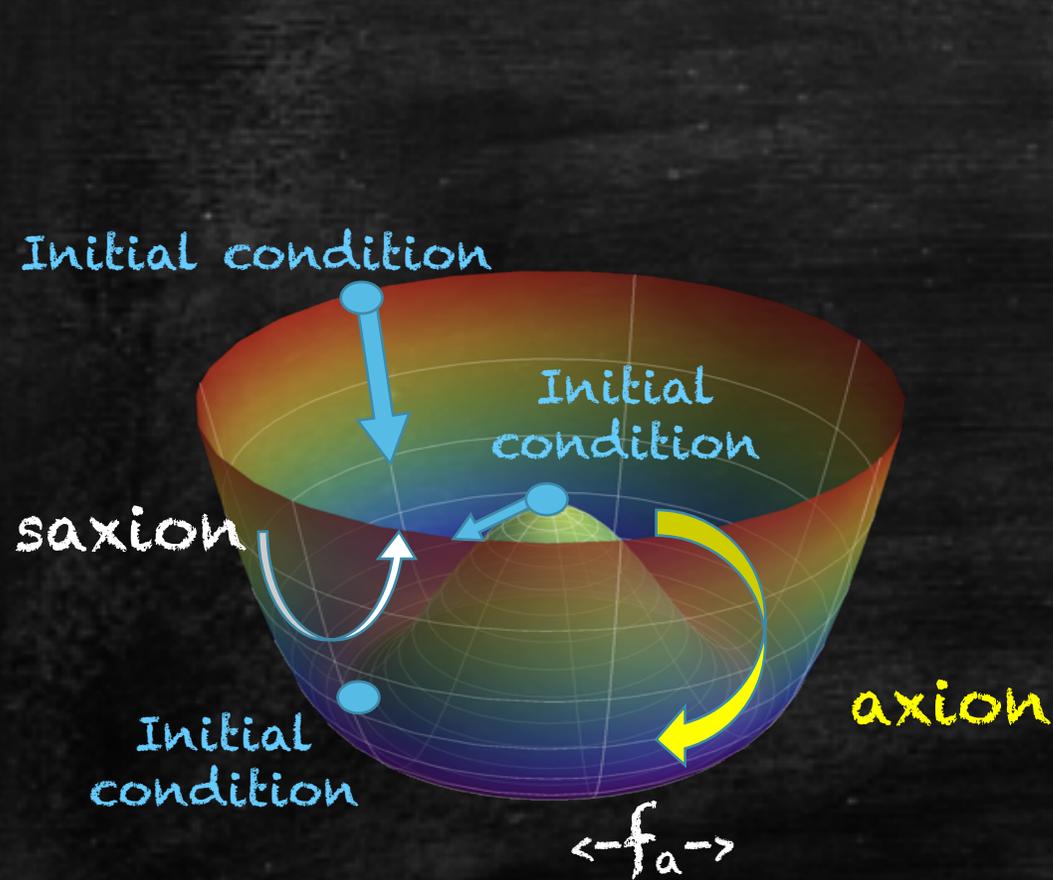
$$\rho_\phi \propto R^{-3(1+w)} = R^{-6}$$

# AXIONS

$$P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$$



# Rotation



# Why Rotation?

Large field value : **Flat potential**

For example, as an initial condition or set dynamically by the Hubble-induced mass

$$V(|P|) \sim -H_I^2 |P|^2 + \frac{|P|^{2d}}{M^{2d-4}}$$

Angular motion : **Explicit PQ breaking**

$$V(P) \sim \frac{P^n}{M^{n-4}} + \text{h.c.}$$

expected from quantum gravity  
or PQ as an accidental symmetry

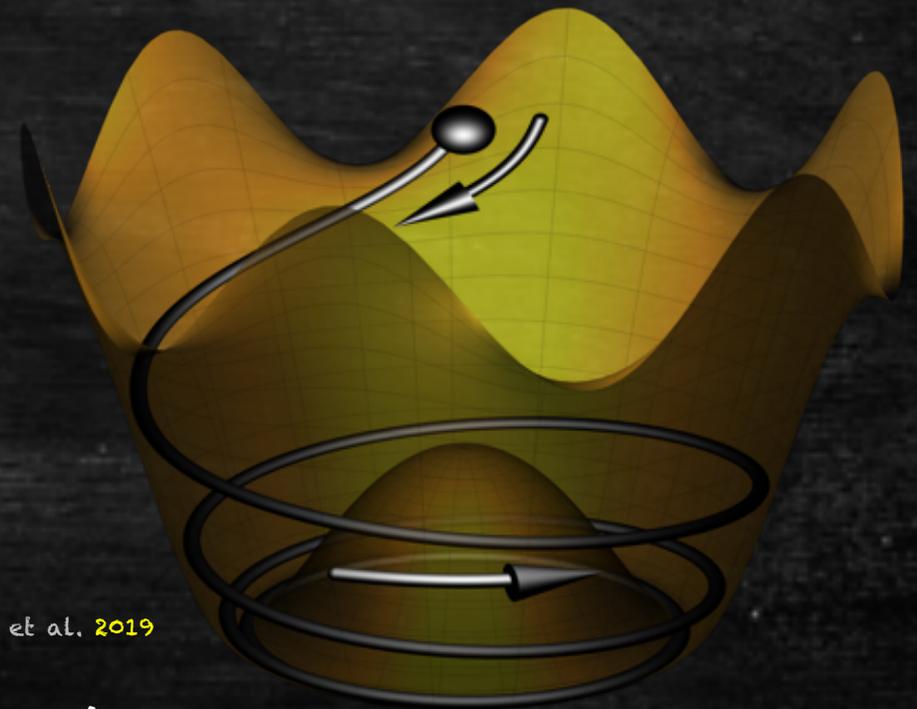
S. Giddings et al. 1988, S. Coleman 1988, G. Gilbert 1988, D. Harlow et al. 2019  
R. Holman 1992, S. Barr 1992, M. Kamiokowski 1992, M. Dine 1992

Dynamics analogous to that in Affleck-Dine baryogenesis

I. Affleck and M. Dine 1991

Initial condition

$$P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$$



PRL 92, 011301 (2004) T. Chiba, F. Takahashi, M. Yamaguchi  
PRL 124, 111602 (2020) RC and K. Harigaya

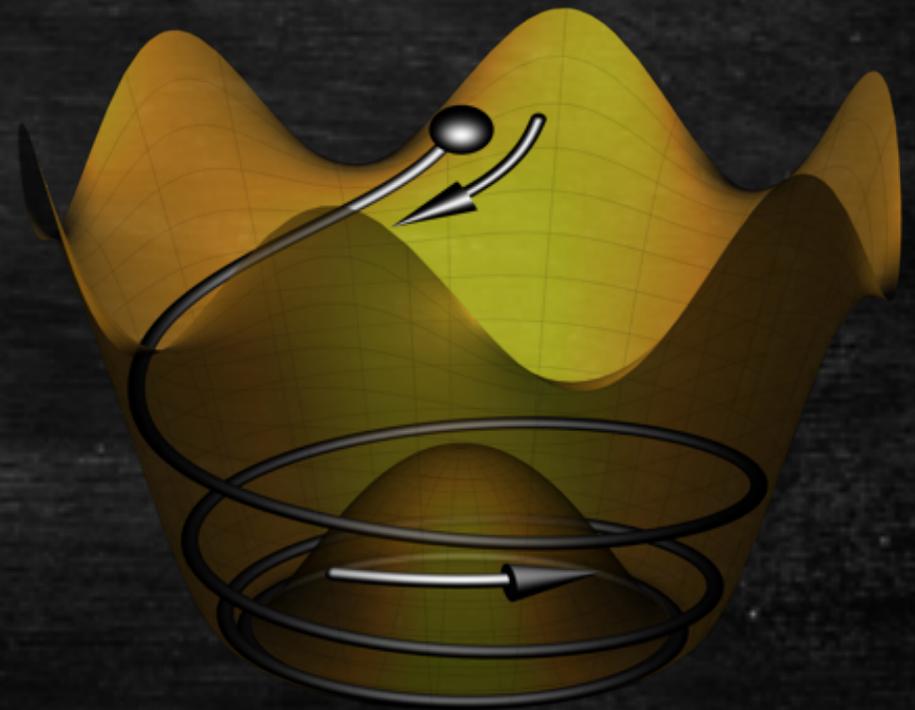
# Asymmetry of PQ Charge

Noether charge associated with the shift symmetry

$$P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$$

$$n_{\text{PQ}} = i P \dot{P}^* - i P^* \dot{P}$$

$$n_{\text{PQ}} = S^2 \dot{\theta}$$



PQ asymmetry  
PQ charge density = Rotation of PQ field

PQ charge is conserved soon after the onset.

# PQ Charge Evolution

Reason:

$$n_{\text{PQ}} = S^2 \dot{\theta} \quad n_{\text{PQ}} R^3 = \text{conserved charge}$$

Conventional:

$$S^2 = f_a^2 \quad \dot{\theta} \propto R^{-3}$$

Our scenario ( $S \gg f_a$ ):

$$\rho_{\text{PQ}} = \dot{\theta}^2 f_a^2 \propto R^{-6}$$

kination!

quartic

$$S^2 \propto R^{-2}$$

$$\dot{\theta} \propto R^{-1}$$

$$\rho_{\text{PQ}} \propto R^{-4}$$

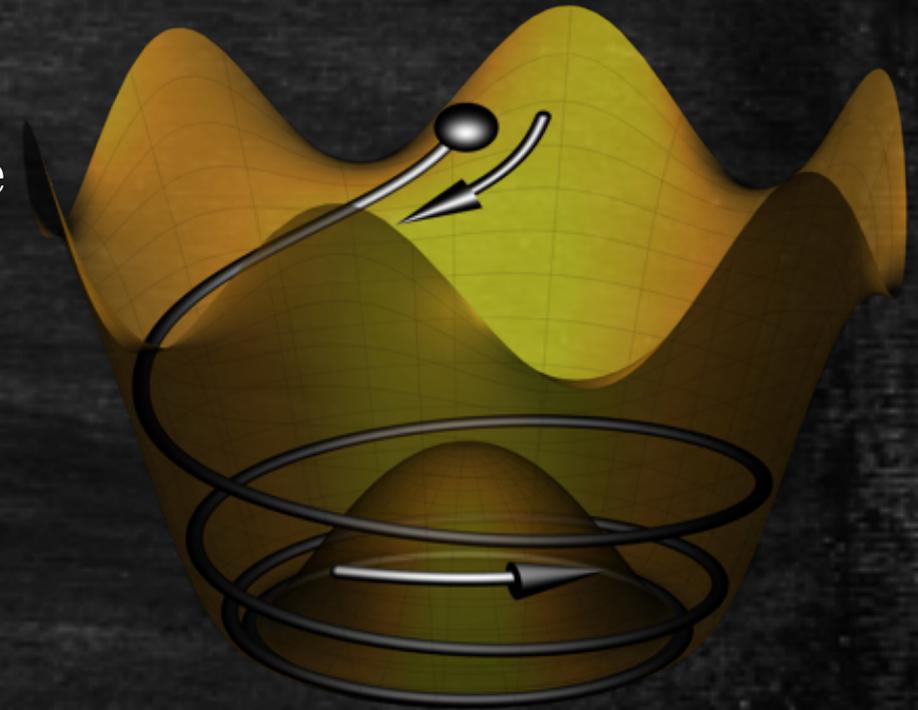
quadratic

$$S^2 \propto R^{-3}$$

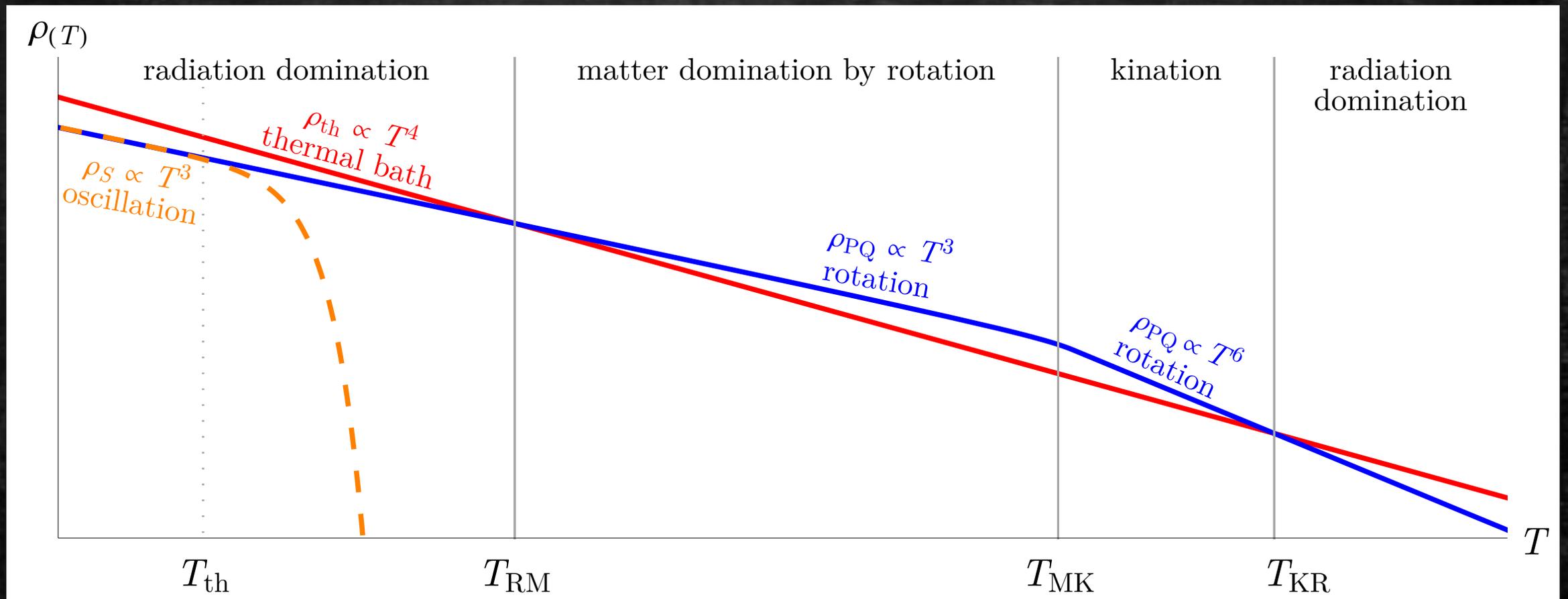
$$\dot{\theta} = \text{constant}$$

$$\rho_{\text{PQ}} \propto R^{-3}$$

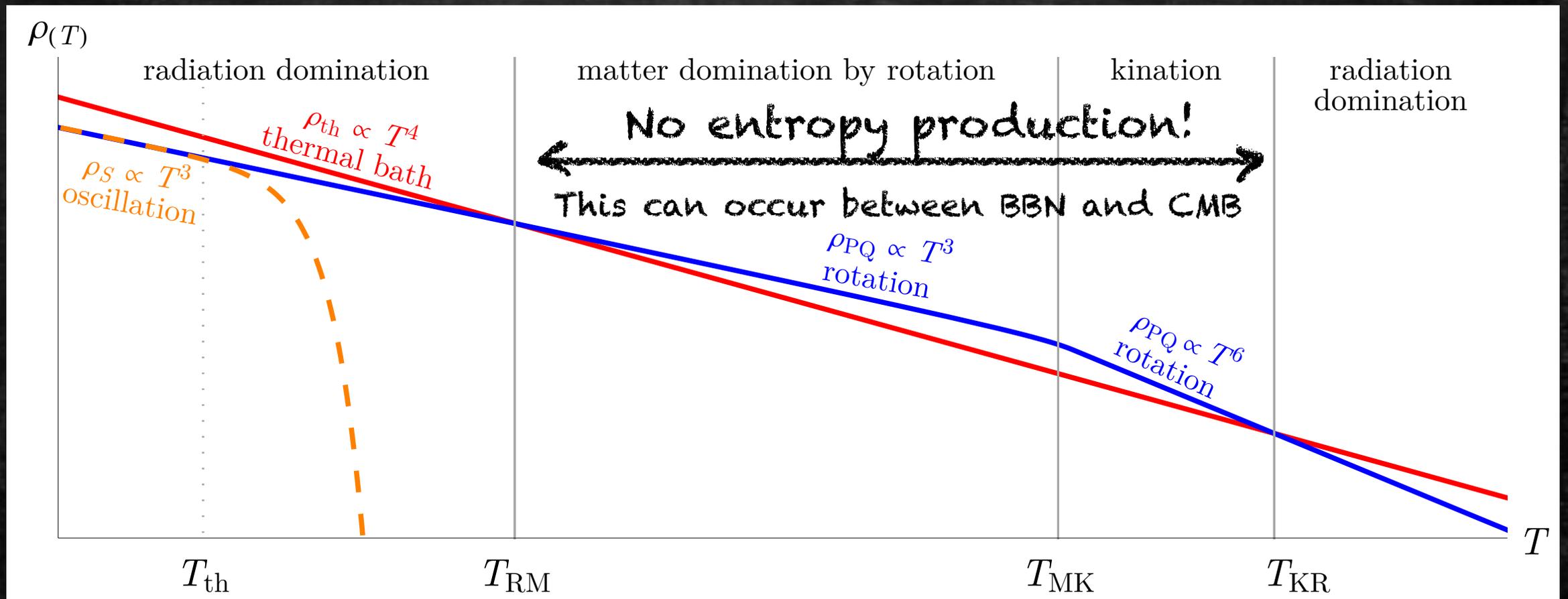
← necessary to achieve kination domination



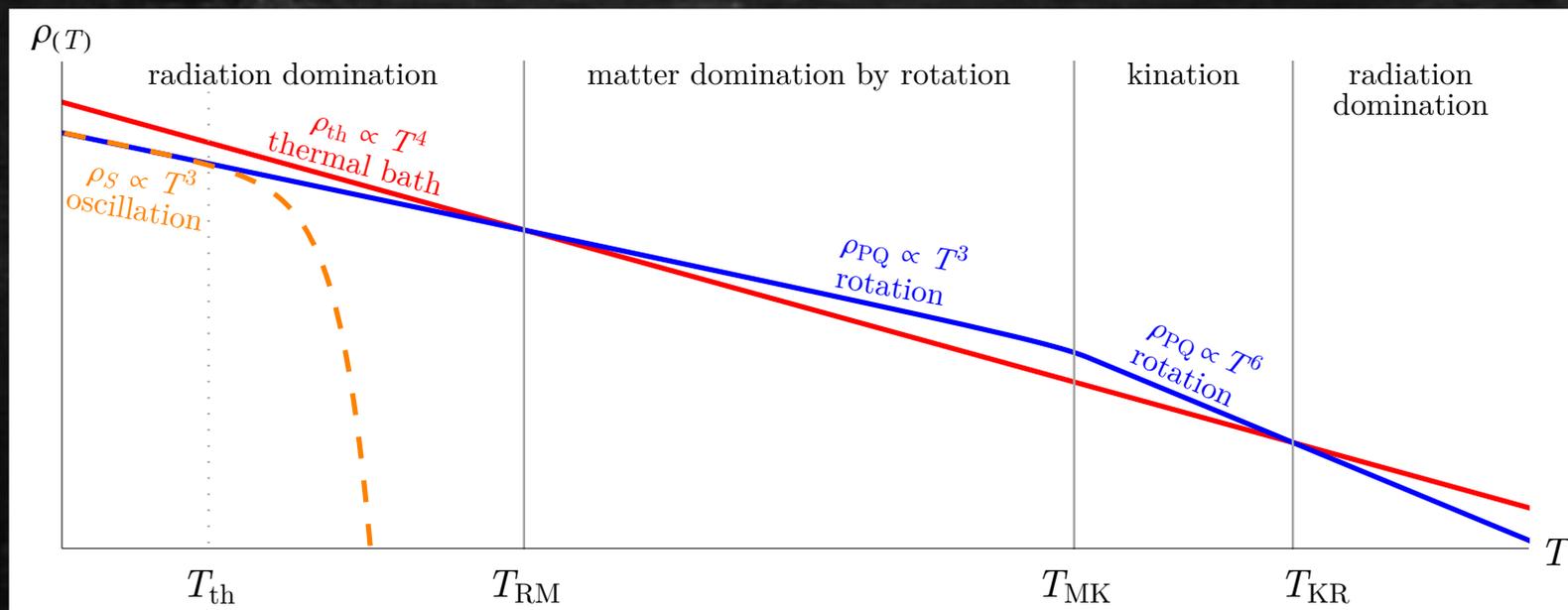
# Evolution of Energy Densities



# Evolution of Energy Densities



# Relevant Temperatures



$$T_{RM} = \frac{4}{3} m_S Y_{PQ}$$

$$T_{MK} = \left( \frac{45}{2\pi^2 g_*} \frac{m_S f_a^2}{Y_{PQ}} \right)^{\frac{1}{3}}$$

$$T_{KR} = \frac{3\sqrt{15}}{2\sqrt{g_*} \pi} \frac{f_a}{Y_{PQ}}$$

$$T_{MK}^3 = T_{KR}^2 T_{RM}$$

# Gravitational Waves from Inflation

The origin is the quantum fluctuations during inflation.

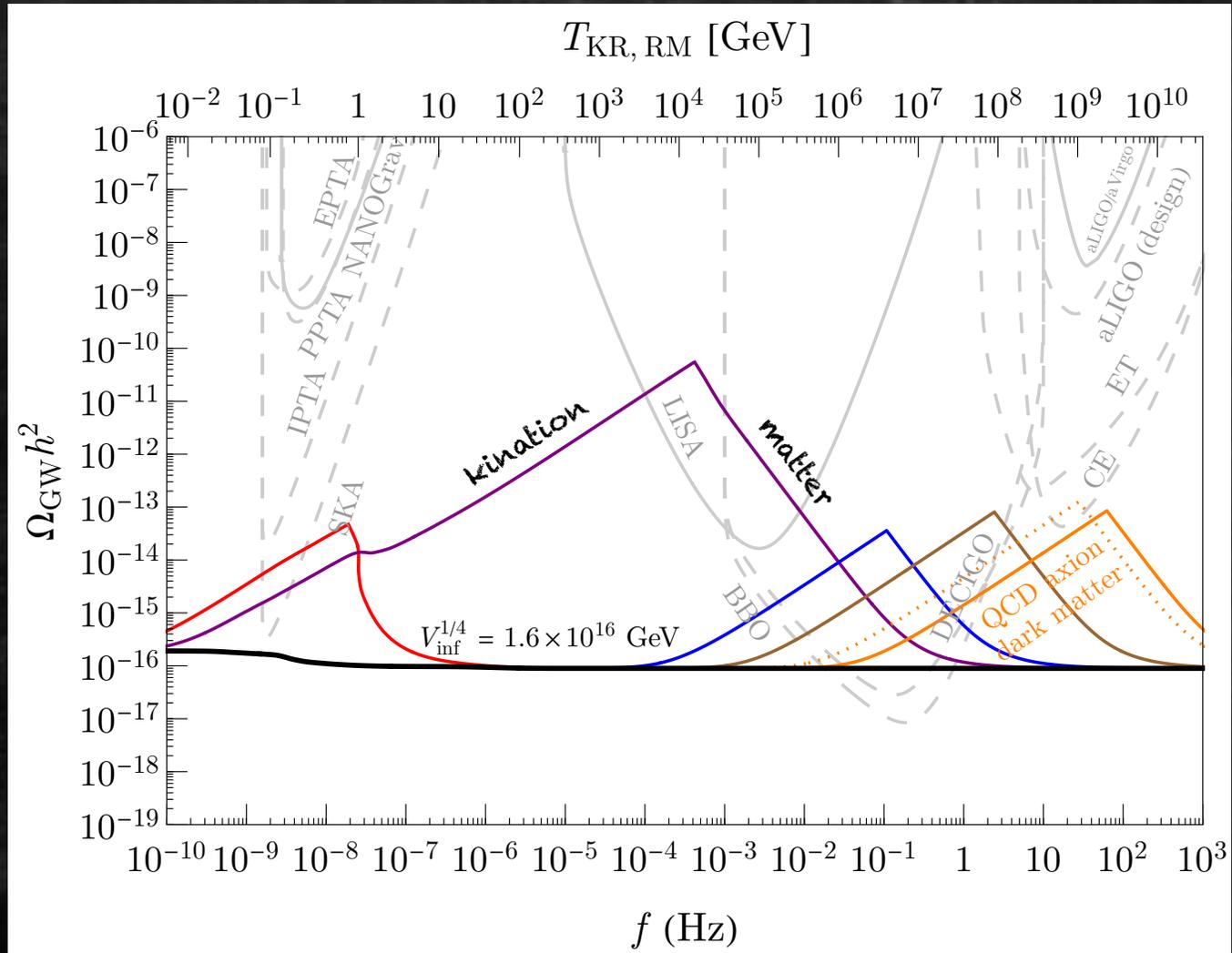
The horizon crossing,  $k = H$ , determines the onset of mode evolutions.

$$\Omega_{\text{GW}} h^2 = \Omega_{r,0} h^2 \frac{1}{\rho_R(T_{\text{hc}})} \frac{k^2}{64\pi G} \mathcal{P}_T(k) \quad \mathcal{P}_T(k) = \frac{2H_I^2}{\pi^2 M_{\text{Pl}}^2}$$

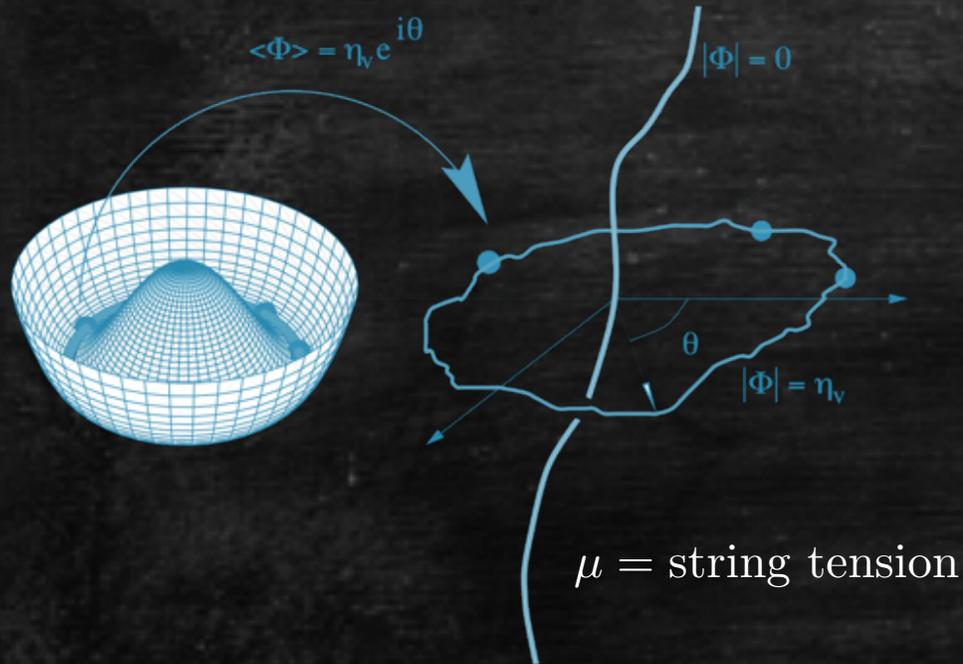
$$k = H(T_{\text{hc}})$$

$$\Omega_{\text{GW}} h^2 \simeq 1.4 \times 10^{-17} \left( \frac{V_{\text{inf}}^{1/4}}{10^{16} \text{ GeV}} \right)^4 \left( \frac{g_{*,\text{SM}}}{g_*(T_{\text{hc}})} \right)^{\frac{1}{3}} \begin{cases} 1 & \text{RD : } f_{\text{RM}} < f \\ \left( \frac{f_{\text{RM}}}{f} \right)^2 & \text{MD : } f_{\text{MK}} < f < f_{\text{RM}} \\ \frac{f}{f_{\text{KR}}} & \text{KD : } f_{\text{KR}} < f < f_{\text{MK}} \\ 1 & \text{RD : } f < f_{\text{KR}} \end{cases}$$

# Gravitational Waves from Inflation kination before BBN



# Gravitational Waves from Cosmic Strings



		Loop formation era		
		Radiation	Matter	Kination
Loop decay era	Radiation	$f^0$	$f^{-1}$	$f^1$
	Matter	$f^{-1/2}$	$f^{-1}$	$f^1$
	Kination	$f^{1/4}$	$f^{-1/2}$	$f^1$

1005.4842 C. Ringeval

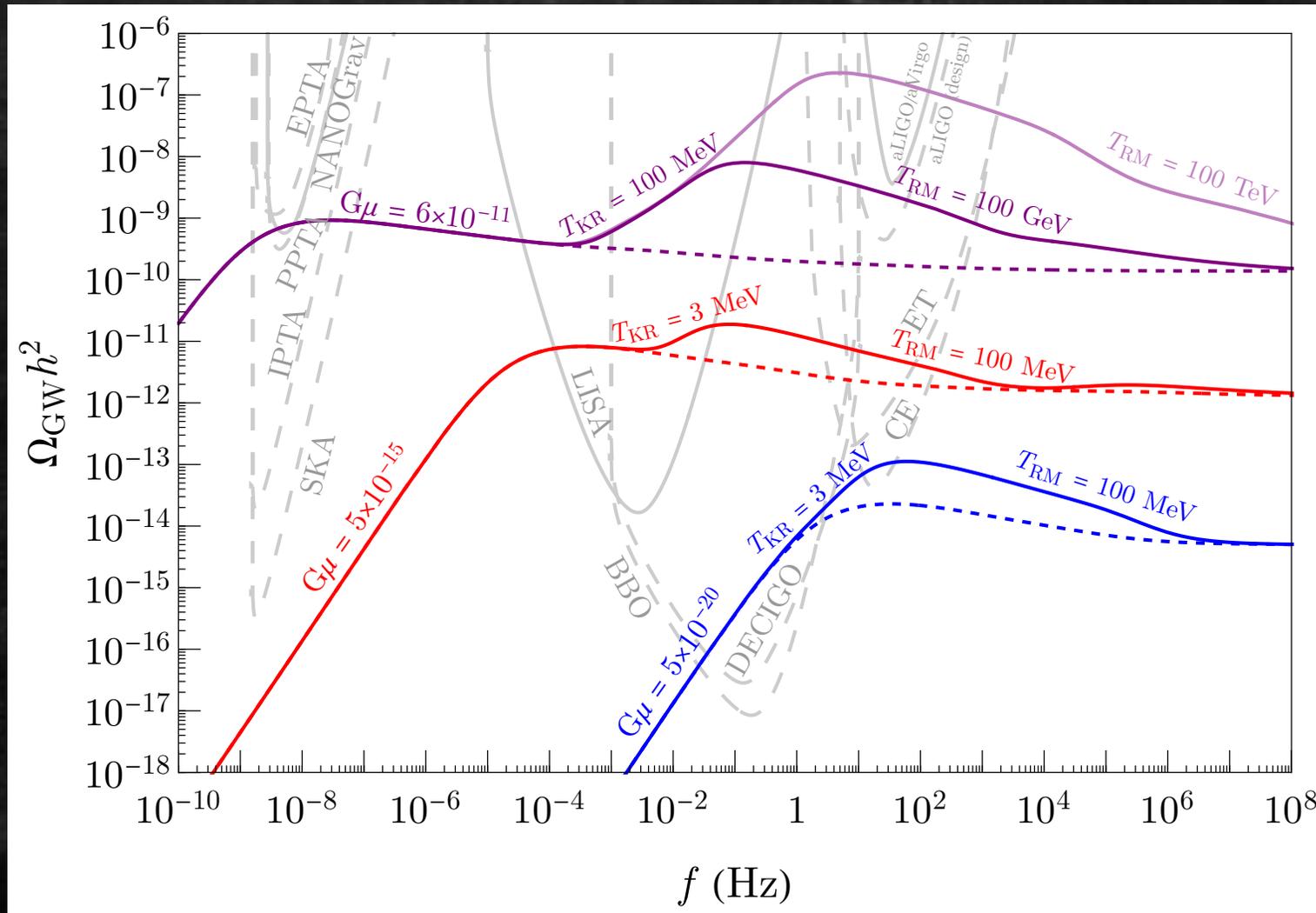


Daniel Dominguez/CERN

1. Flat spectrum for radiation domination
2. An enhanced Hubble increases loop energy density.

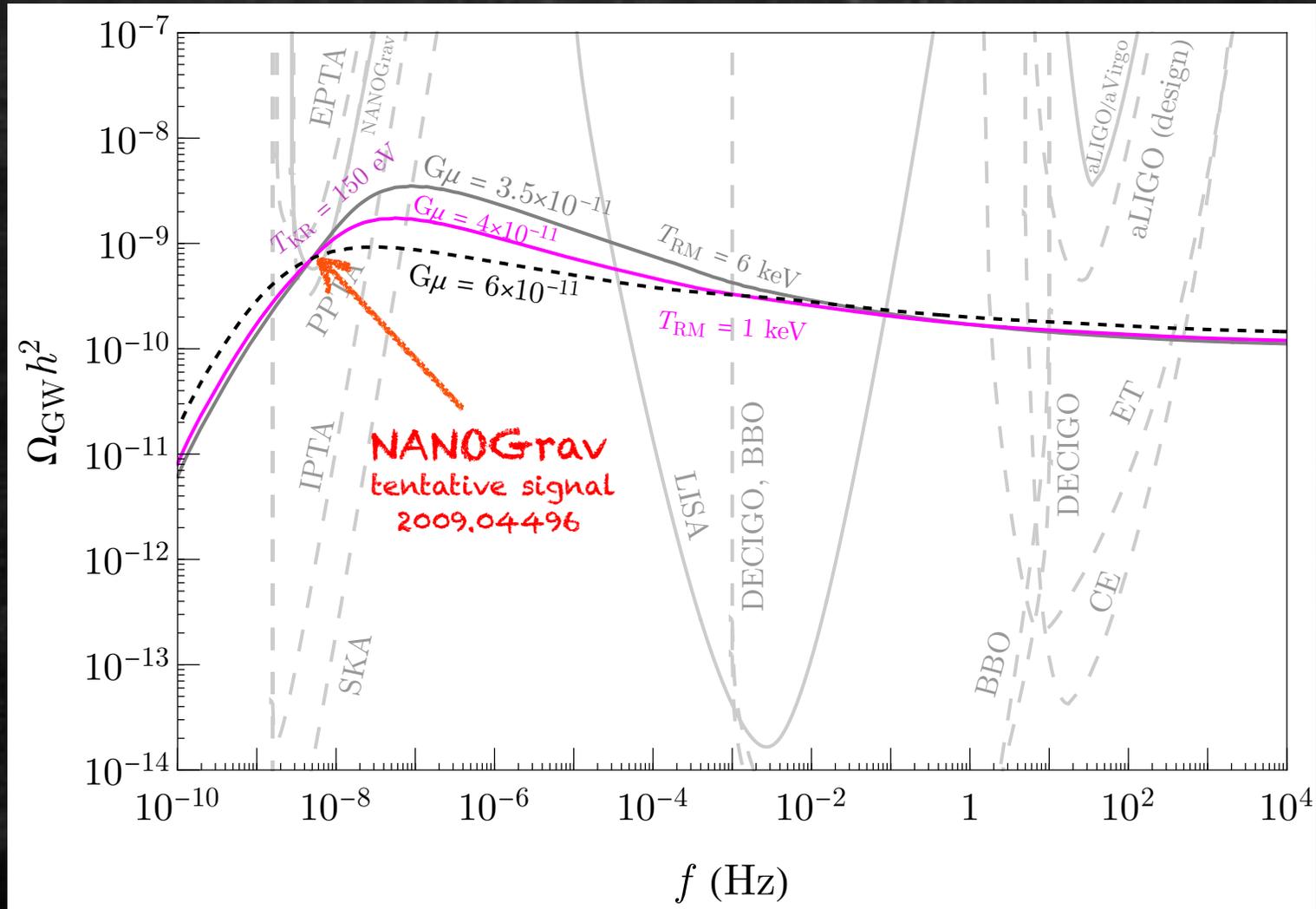
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kination before BBN



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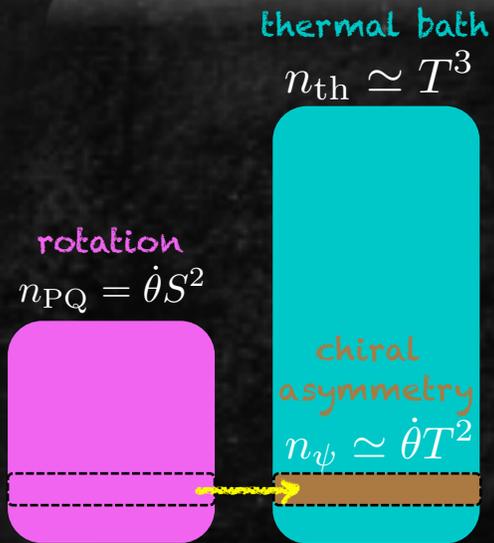
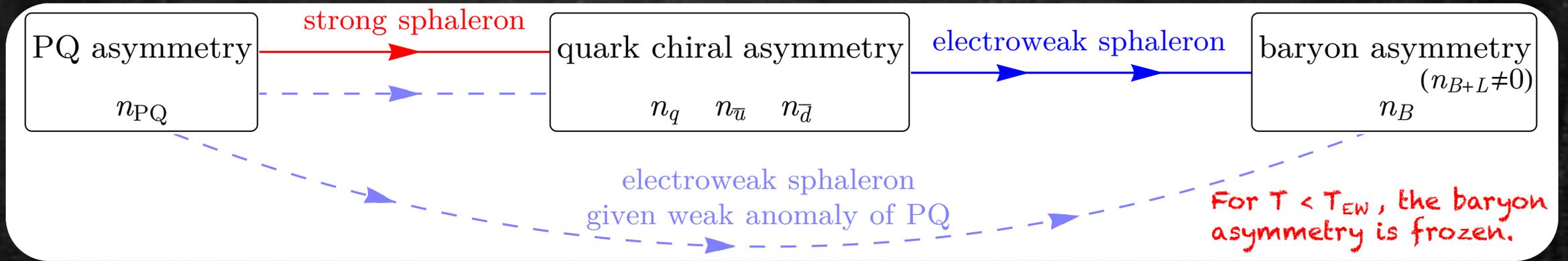
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# Axiogenesis

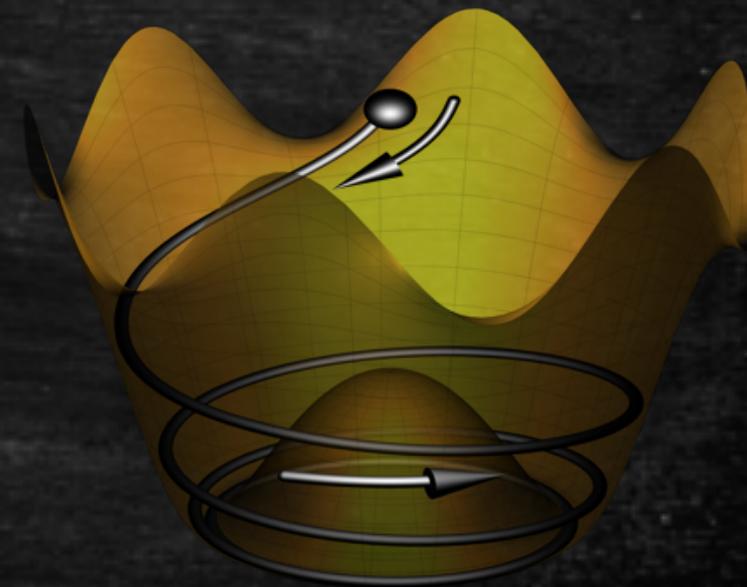
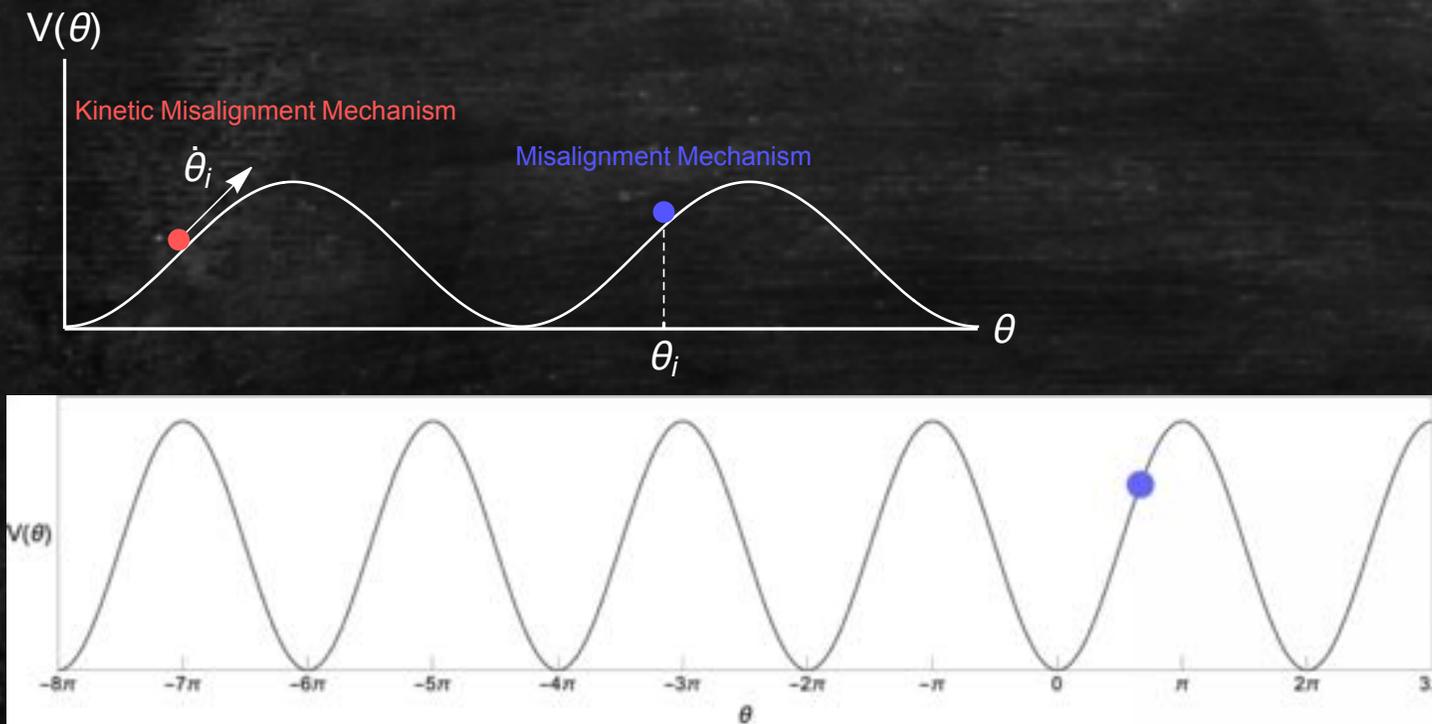


$$Y_B \equiv \frac{n_B}{s} = \frac{c_B \dot{\theta} T^2}{s} \Big|_{T=T_{EW}} = c_B Y_{PQ} \left( \frac{T_{EW}}{f_a} \right)^2$$

Baryon asymmetry fixes rotational speed, equivalently  $Y_{PQ}$ .

# Kinetic Misalignment Mechanism

"an alternative scenario where the axion field has a nonzero initial velocity"



Consequence: delaying\* usual  $T_{osc}$  until  $KE = PE$ , enhancing the dark matter abundance

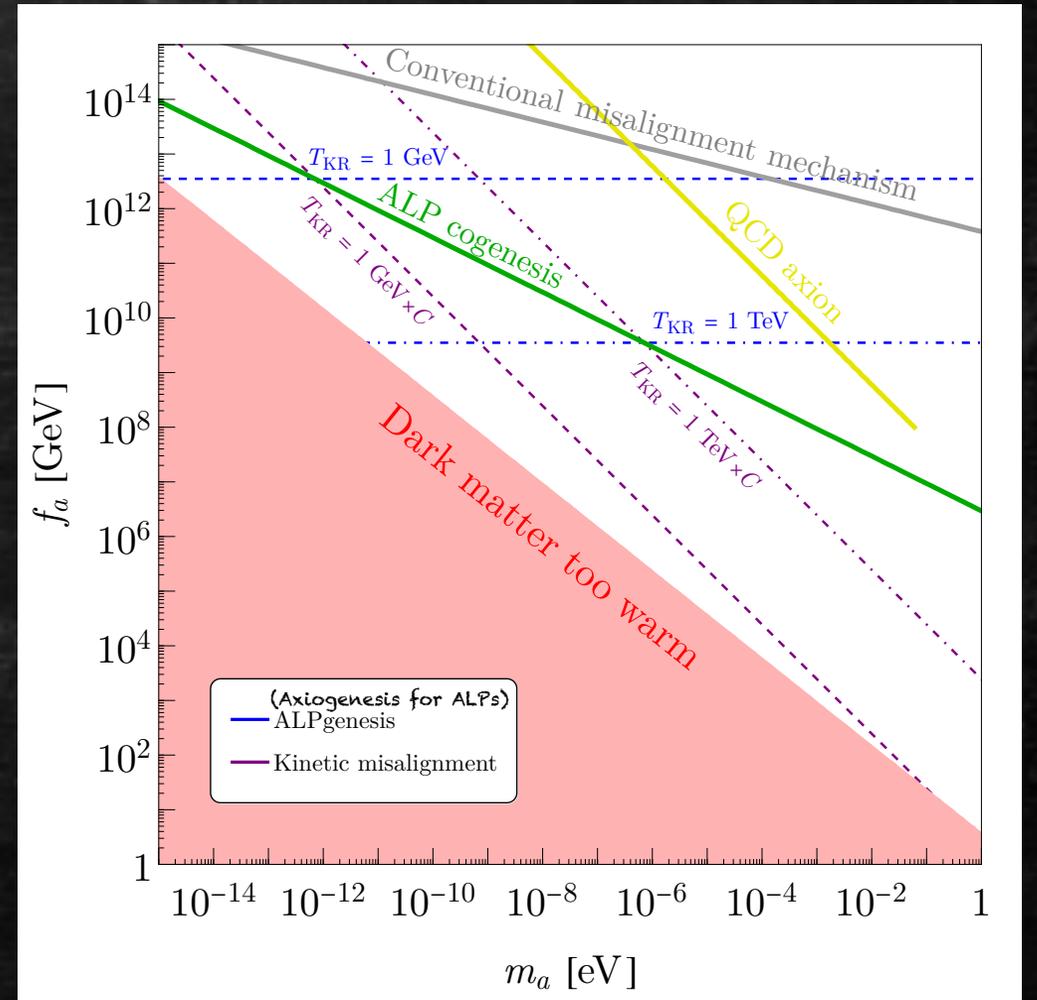
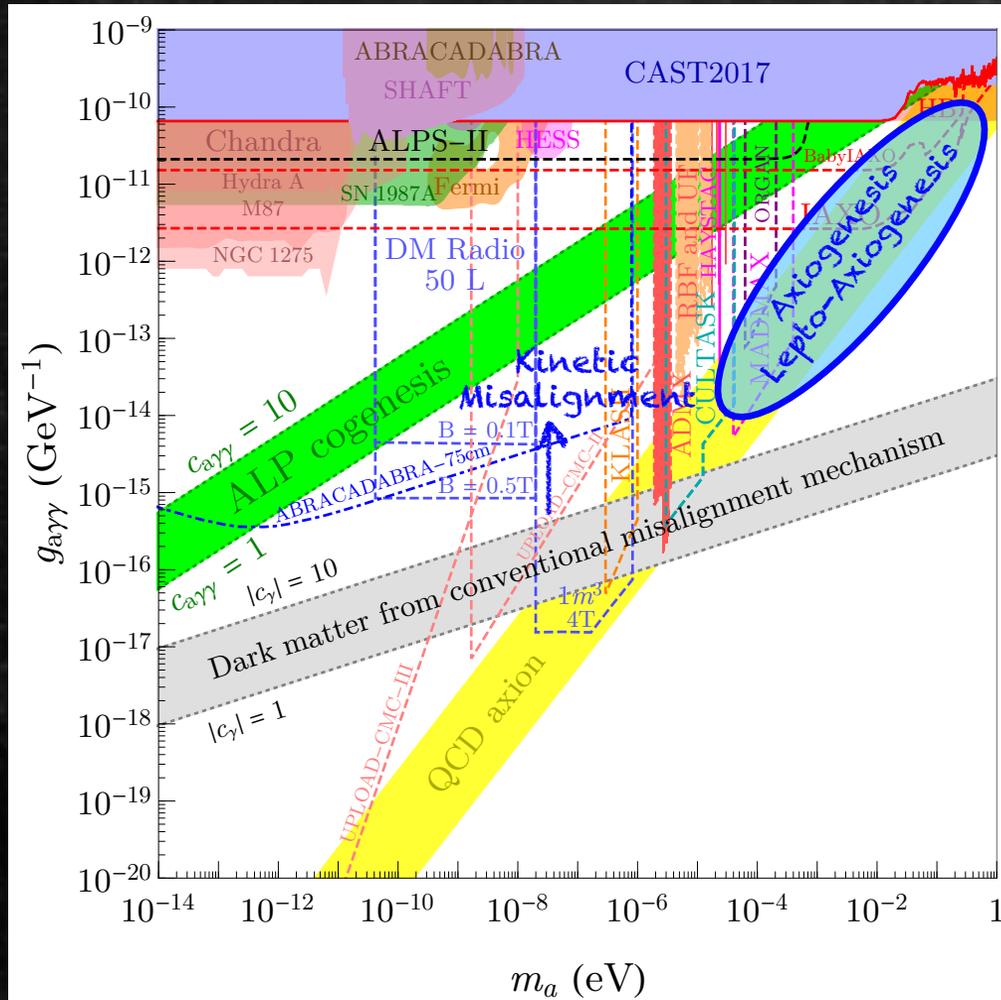
$$\text{Abundance: } \frac{\rho_a}{S} = C m_a Y_{PQ}$$

\* Parametric resonance in fact occurs before oscillations start. The abundance is modified by an  $O(1)$  factor but axion dark matter can be warm.

arXiv:2104.02077 RC, K. Harigaya, A. Pierce

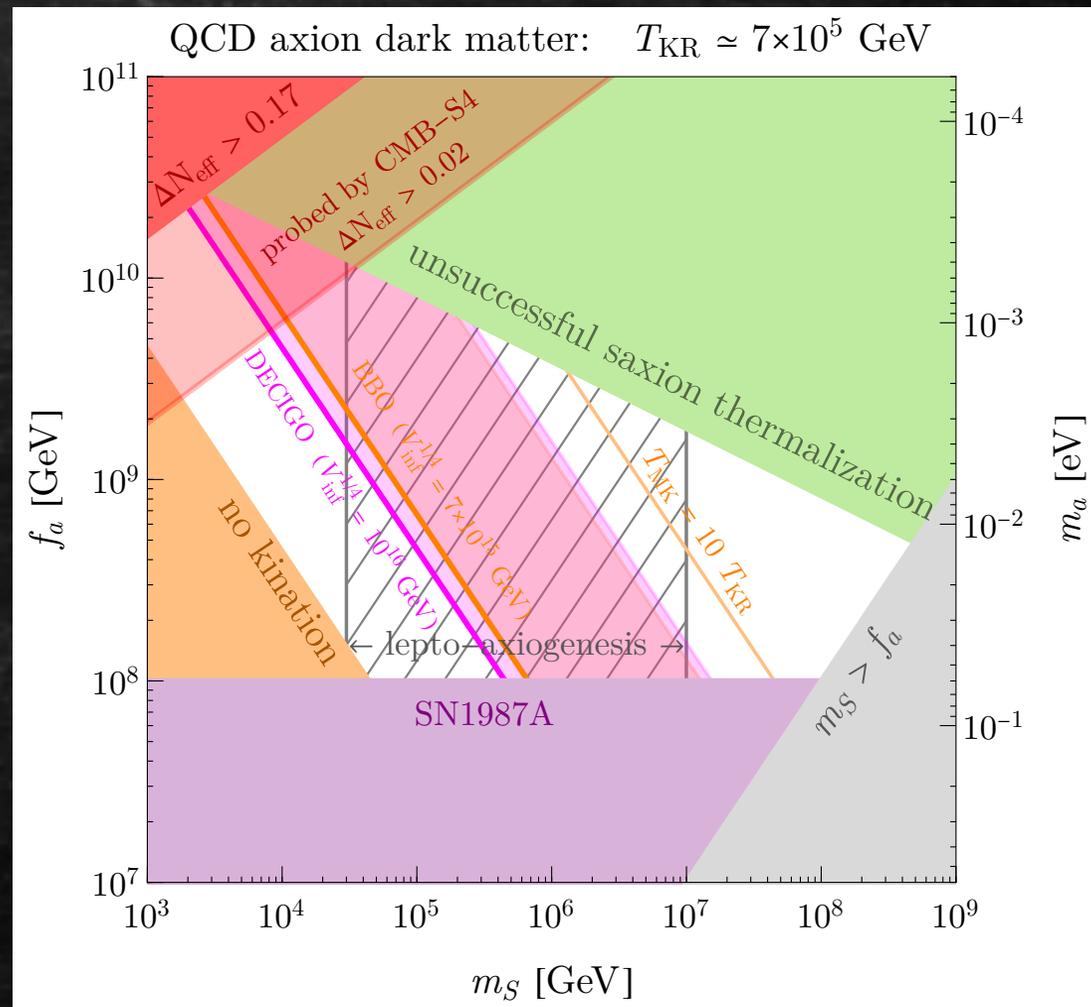
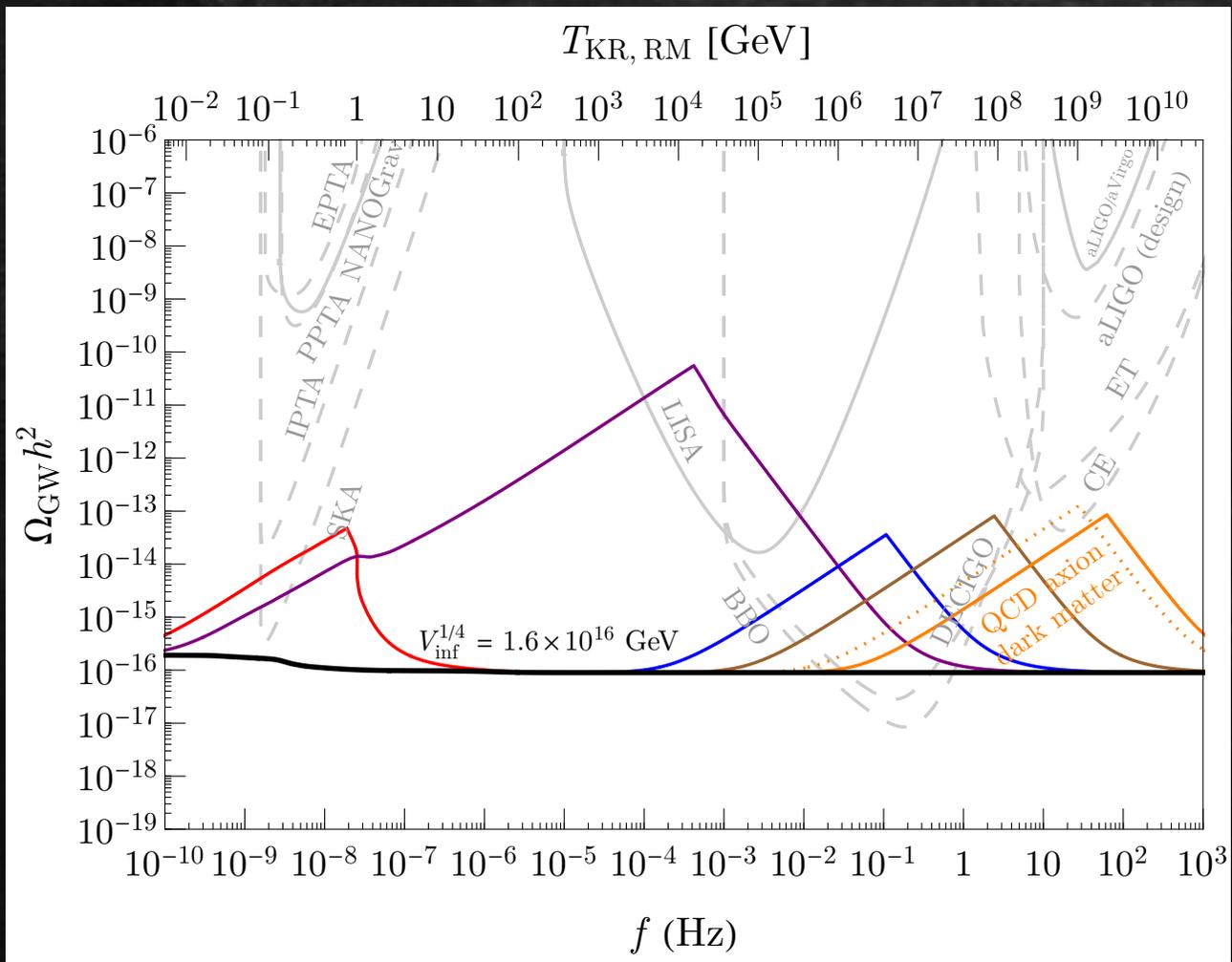
PRL 124, 251802 (2020) RC, L. Hall, K. Harigaya

# Predictions from Kinetic Misalignment and Axionogenesis



# The QCD Axion

probed by gravitational waves from inflation



# Probing PQ-breaking Potential

Piecewise approximation

$$\rho_\theta \propto \begin{cases} a^{-3} & \text{for } S \gg f_a \text{ i.e. } T \gg T_{\text{MK}} \\ a^{-6} & \text{for } S \simeq f_a \text{ i.e. } T \ll T_{\text{MK}} \end{cases}$$

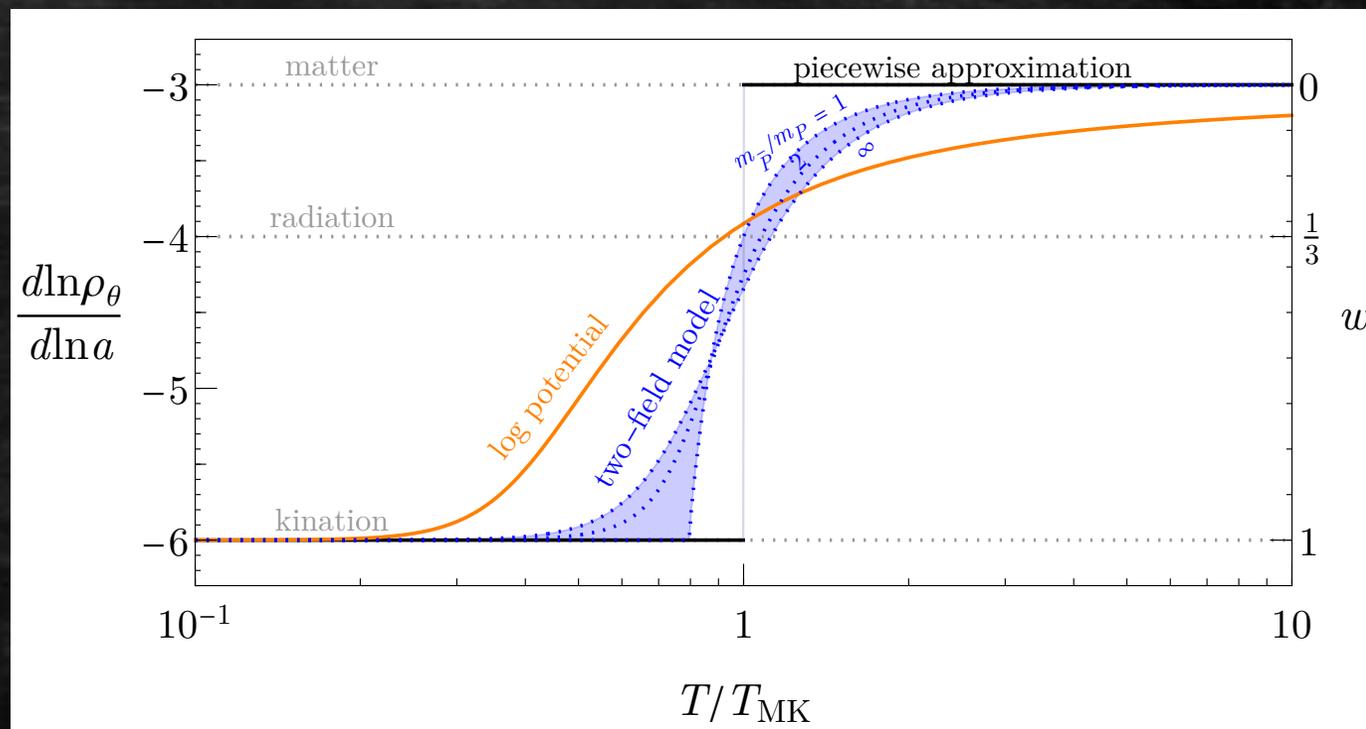
Log potential

$$V(P) = m_S^2 |P|^2 \left( \ln \frac{2|P|^2}{f_a^2} - 1 \right)$$

Two-field model

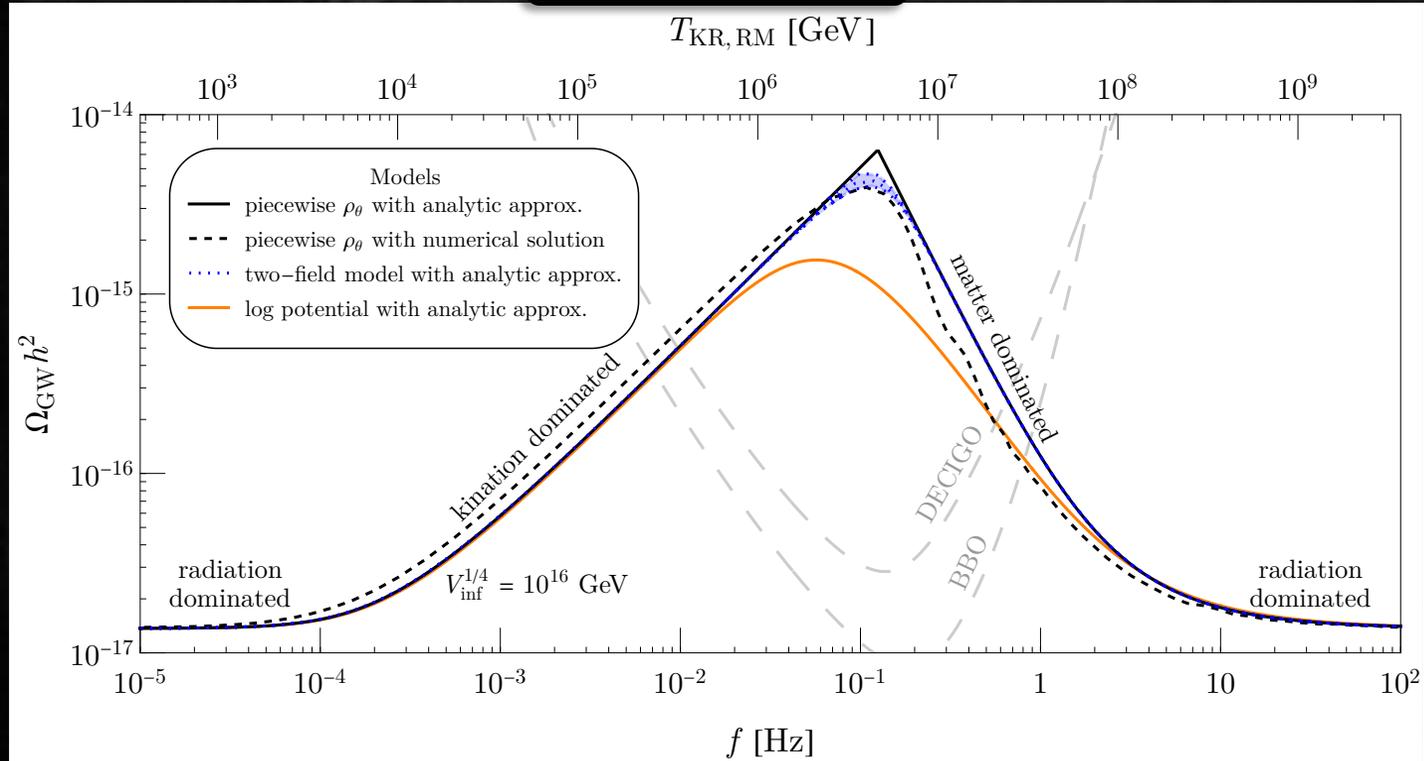
$$W = X(P\bar{P} - v_P^2)$$

$$V_{\text{soft}} = m_P^2 |P|^2 + m_{\bar{P}}^2 |\bar{P}|^2$$

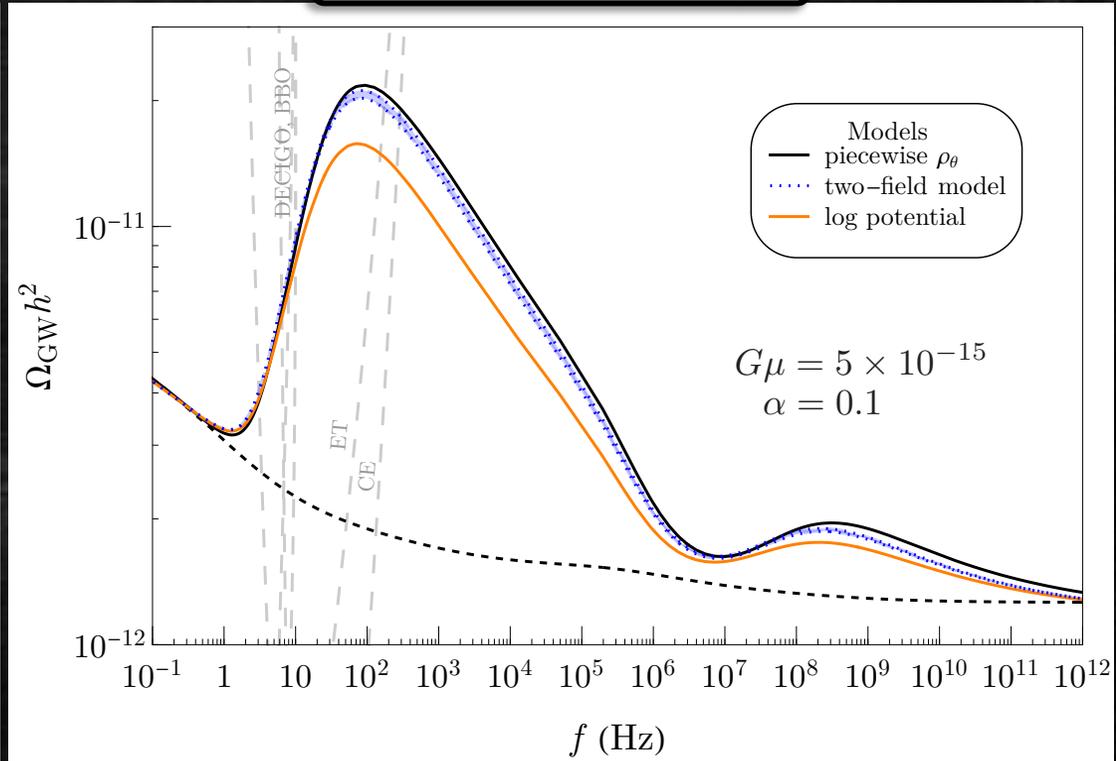


# Probing PQ-breaking Potential

from inflation

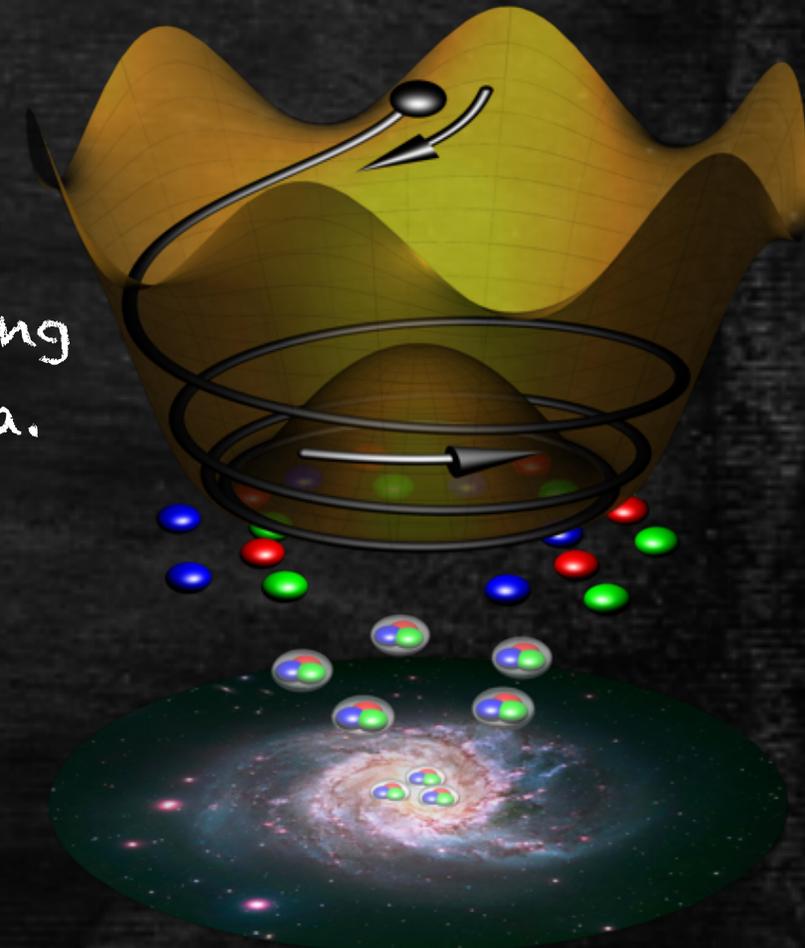


from cosmic strings



# Conclusions

- ✓ **New axion dynamics** allows the QCD axion to simultaneously explain
  - ✓ the Strong CP problem
  - ✓ the dark matter abundance
  - ✓ the baryon asymmetry
- ✓ This paradigm predicts axion kination, featuring a triangular peak in gravitational wave spectra.
- ✓ Other possible signatures:
  - ✓ (QCD) axion searches
  - ✓ Warm axion dark matter
  - ✓ Matter power spectrum
- ✓ New model building opportunities



Thank you!

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RIP, Norman.